

REACH

Scientific Overview of Extremely High-Contrast Spectroscopy at the Subaru Telescope

Hajime Kawahara (U Tokyo)
REACH collaboration

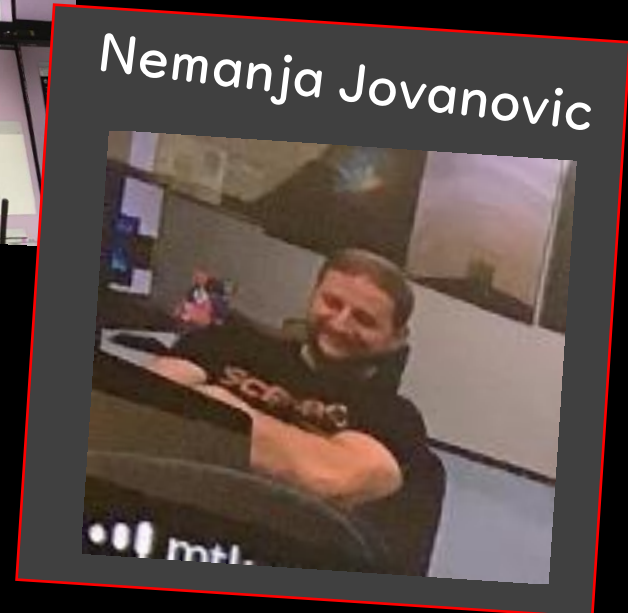
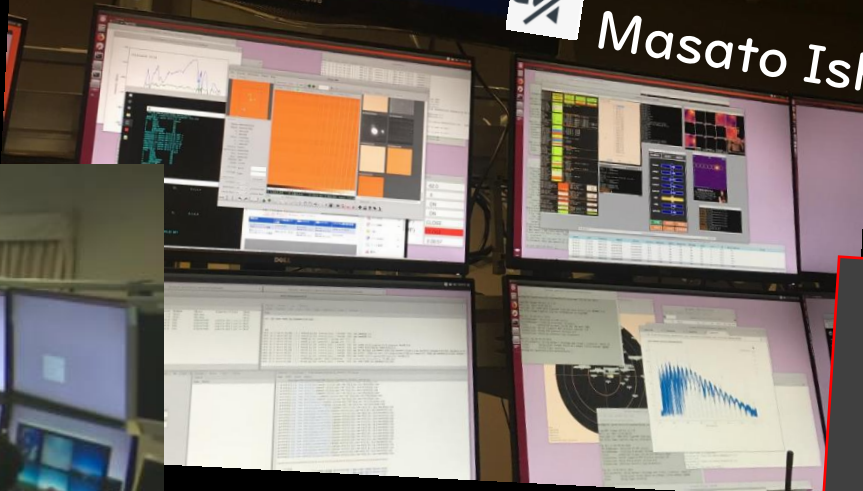


Takayuki Kotani


Ananya Sahoo



Masato Ishizuka



Nemanja Jovanovic

The first real  spectrum by REACH in Oct 16th (2019)

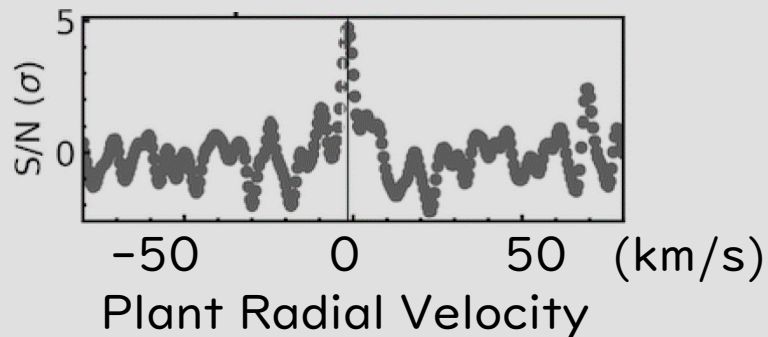
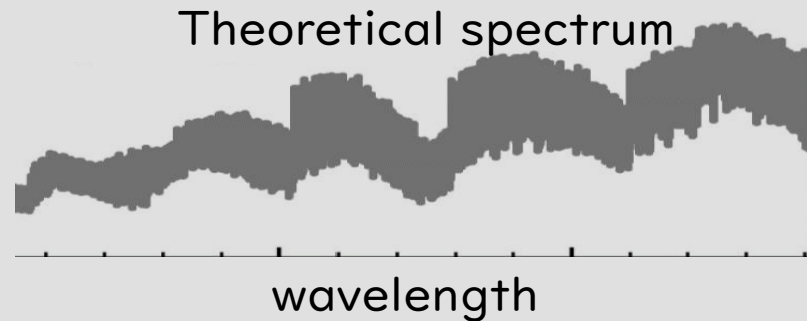
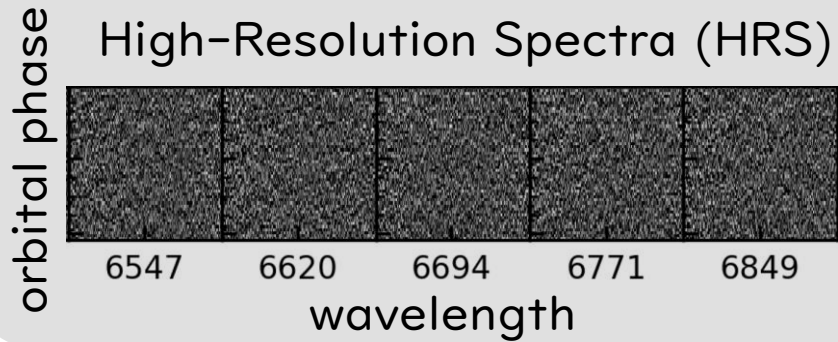


Sebastien Vievard

Olivier Guyon

Julien Lozi

HRS detection of molecules in planetary atmosphere



Transmission CO (Snellen+2010), Metals (Hoeijmakers+)

Emission CO (Brogi+2012 etc.), H₂O (Birkby+2013)

TiO (Nugroho, *H.K.*+2017), HCN (Cabot+2019)

Imaging CO, H₂O (Konopacky+2013, Snellen+2014, Hoeijmakers+)

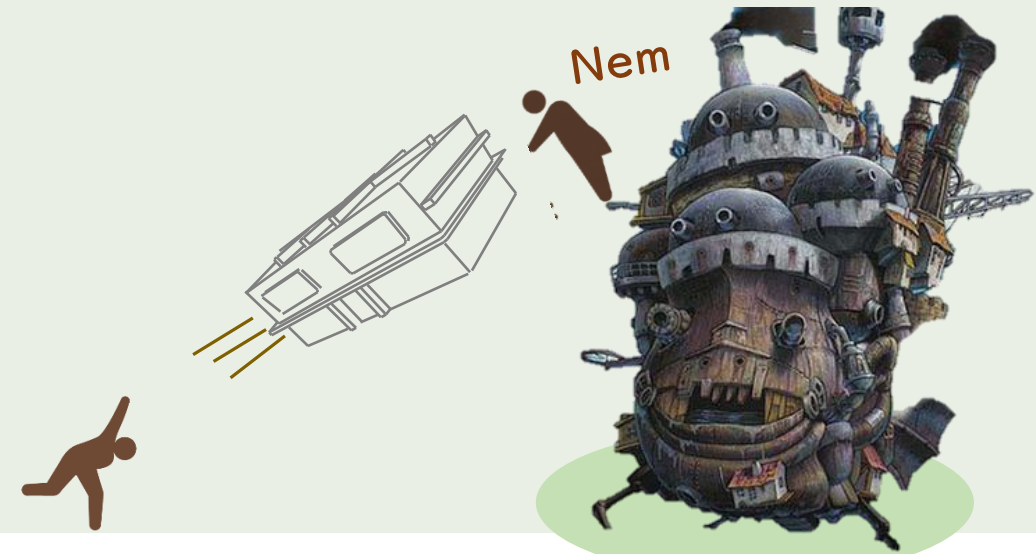
HRS with Coronagraph? No scientific results yet

e.g) Kawahara, Murakami, Matsuo, Kotani, *ApJS* 212, 27 (2014),

“Spectroscopic Coronagraphy for Planetary Radial Velocimetry of Exoplanets”

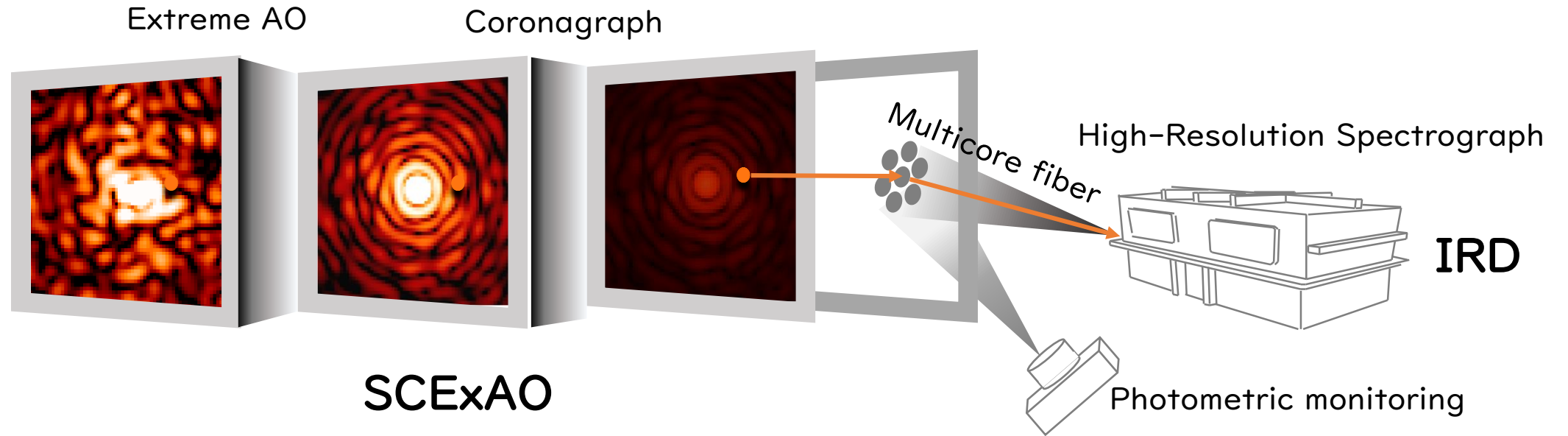
Snellen+2015, Jovanovic+2017, Wang+2017, also Sparks & Ford 02

REACH: HCI (SCEXAO) + HRS (IRD) since 2014





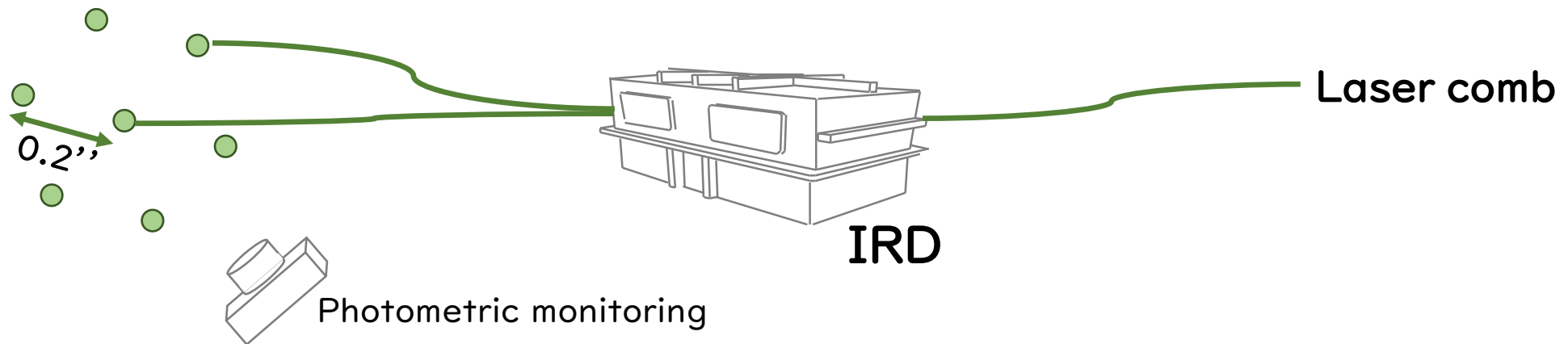
High Dispersion Coronagraph (HDC or HRS-C) by REACH project



Recipe of REACH

A Select your favorite (or no) coronagraph (Lyot, 8oct, vortex so on) in SCExAO

B The REACH multicore fiber has 7 ports. Select two inputs to IRD.



C Available Y, J, and H with $R = 100,000$ thanks to SMF

Advanced option 1: You can use 1 in 2 fibers for laser comb.

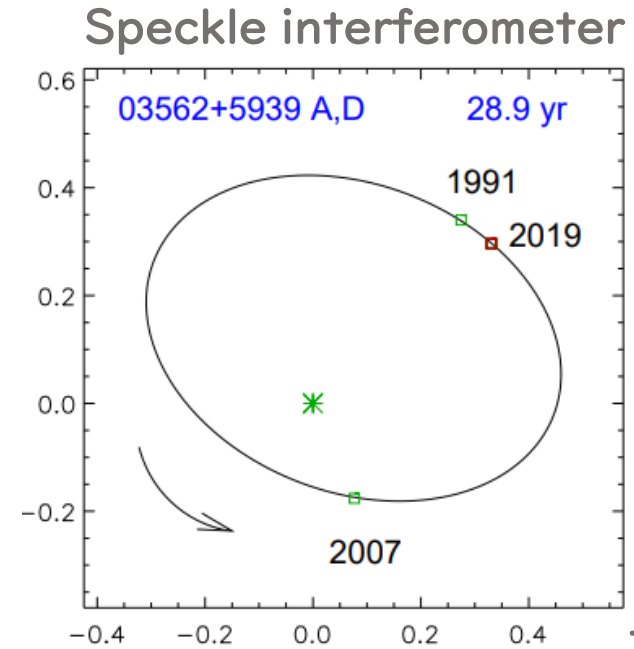
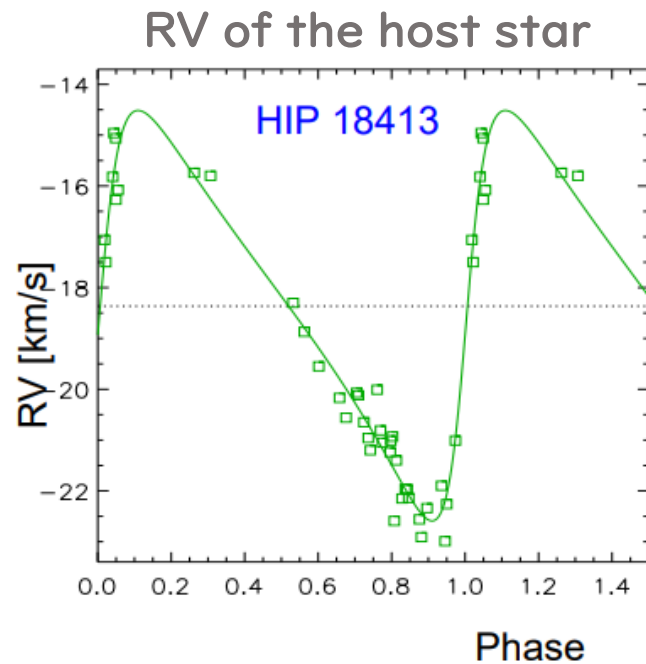
Advanced option 2: The REACH monitoring camera for speckle nulling.



Demonstration



Target: Spectroscopic Binary (SBI) HIP 18413



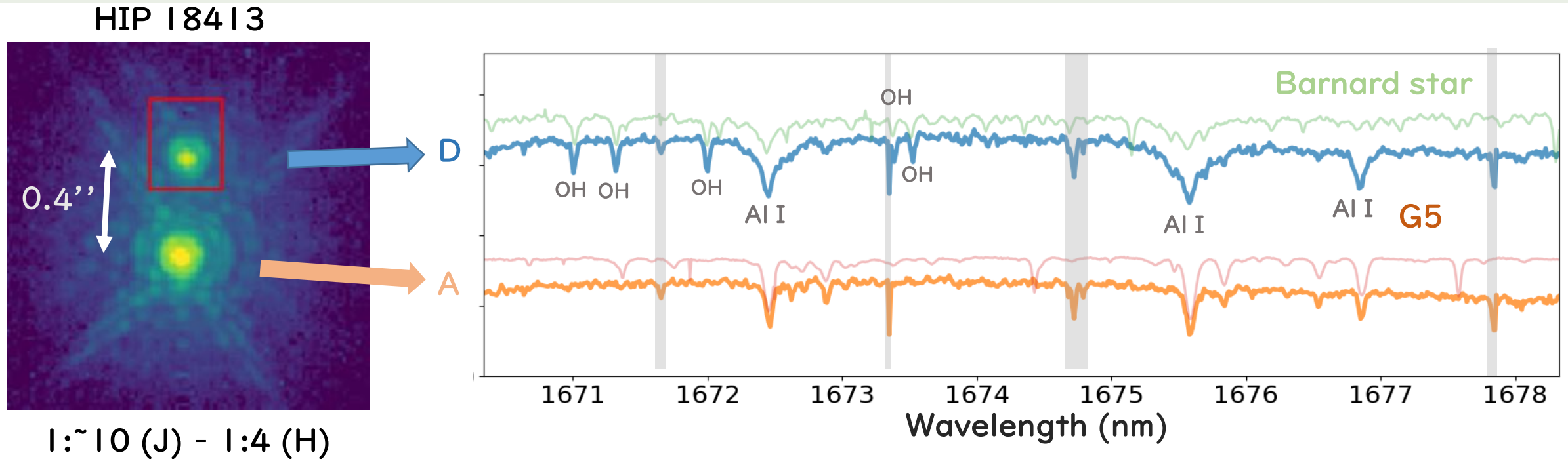
The lines from a companion was unavailable (SBI).



Demonstration



HRS with No Coronagraph by REACH



We got a spectrum of a companion D, completely separated from A!
REACH reveals HIP 18413B is early-mid M.

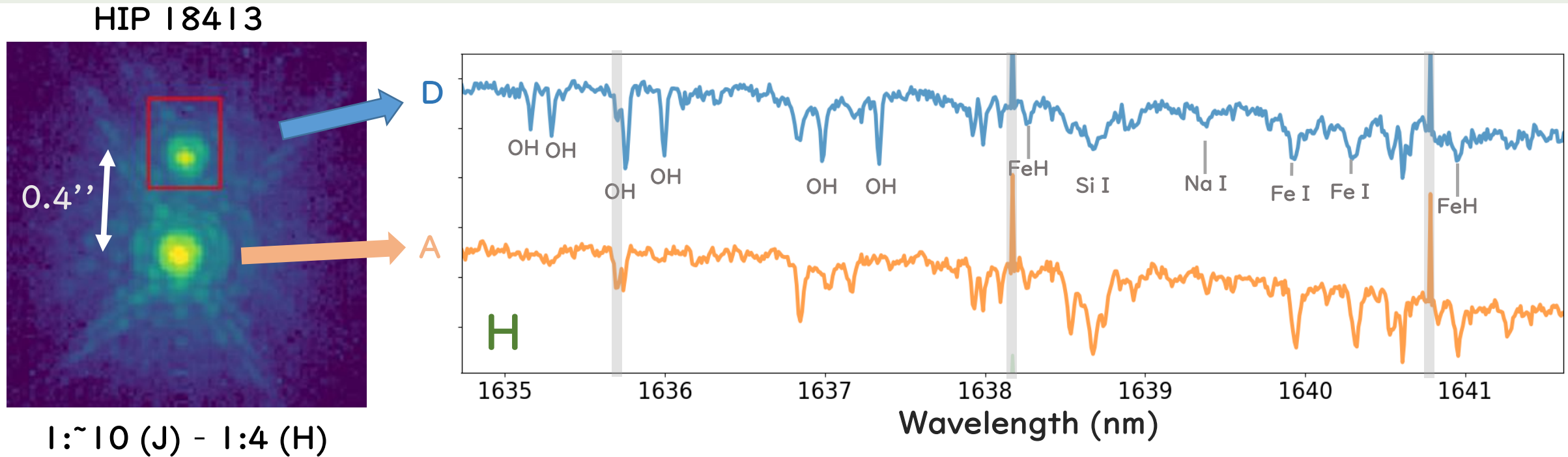
From engineering run in Oct 16th (2019)



Demonstration



High S/N spectrum in 10 min exposure for a M-type companion at 23 pc



Relative RV ~ 0 km/s \gt more precise CCF analysis is needed \gt dynamical masses of the stars

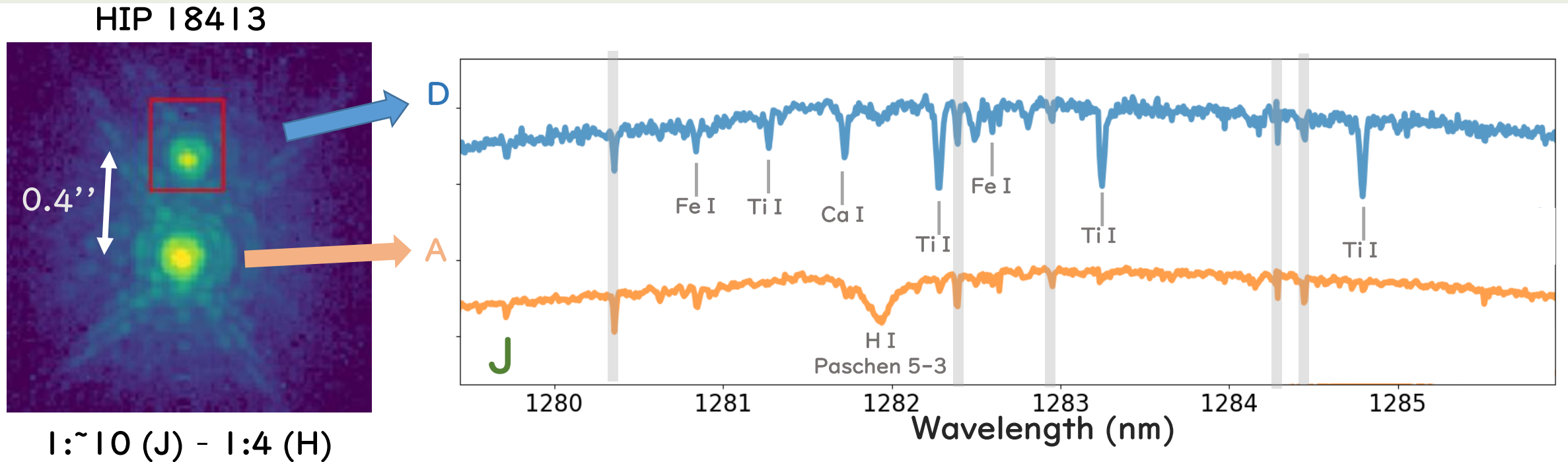
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Demonstration



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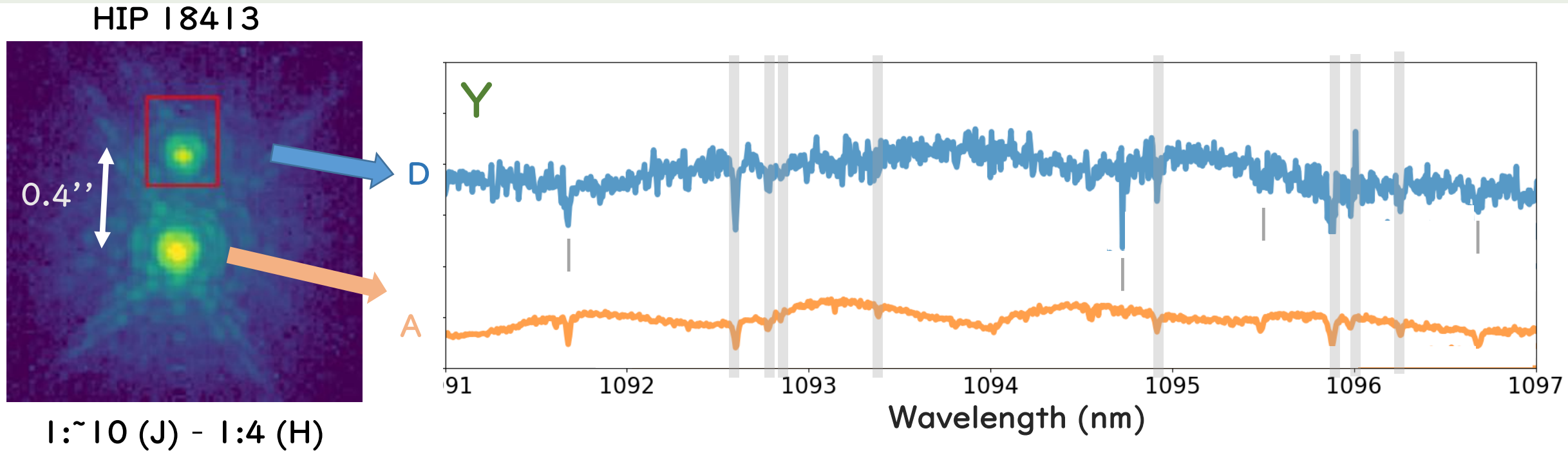
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Demonstration



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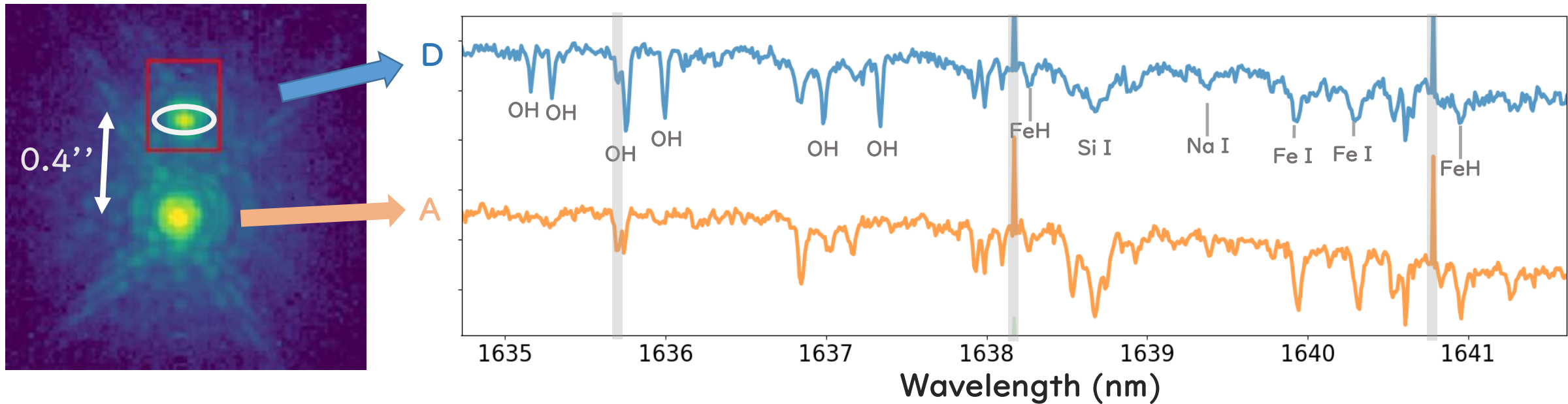


Relative RV ~ 0 km/s \gt more precise CCF analysis is needed \gt dynamical masses of the stars

From engineering run in Oct 16th (2019)

Science case # I: Precise Radial Velocimetry of a Companion Star

REACH can reach hidden nearby M-type targets for precise radial velocity survey with laser frequency comb



- Planet around a binary companion
 - Dynamical mass of a companion
 - Dynamical planet mass

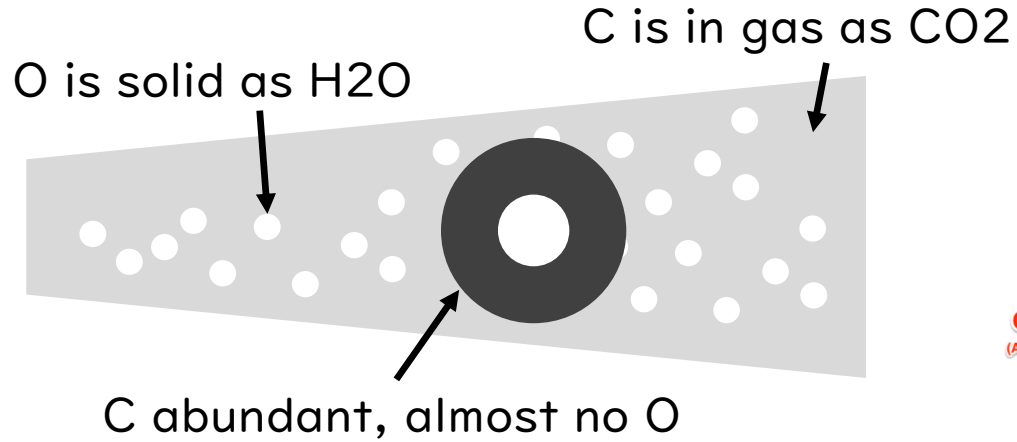
- Planet in the HZ around a companion
 - Complex climate?
 - New survey space



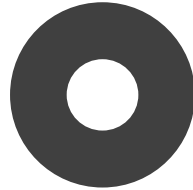
Science case #2: HRS-C for cool directly imaged planets



- In situ formation of gas giants



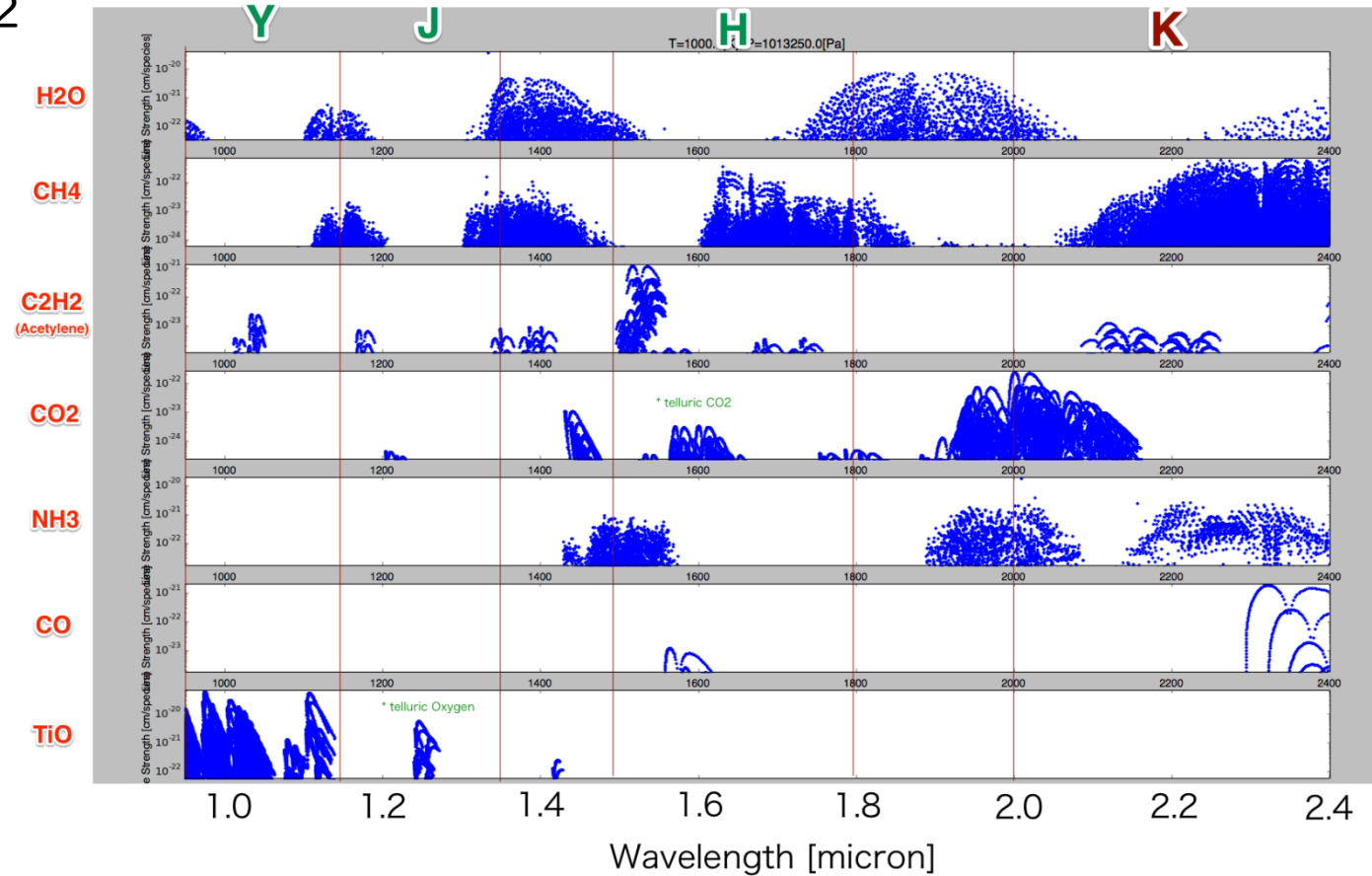
↓ evolution



CH₄ rich atmosphere
Current Jupiter?

HCN is also available in H band

Use KPIC!

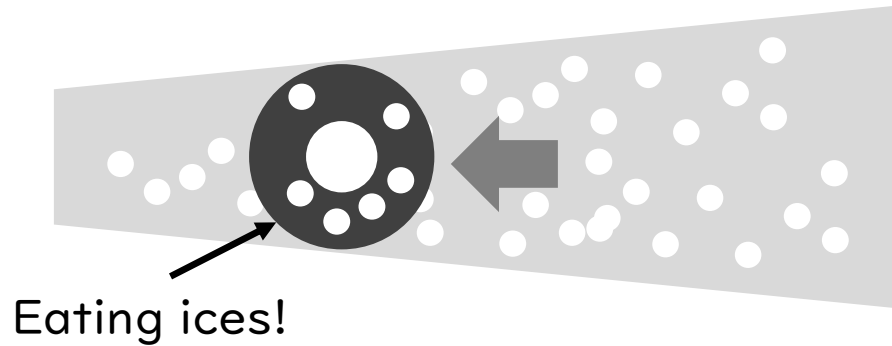




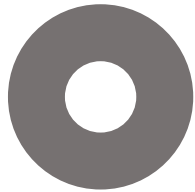
Science case #2: HRS-C for cool directly imaged planets



● Migration injects oxygen in atmosphere



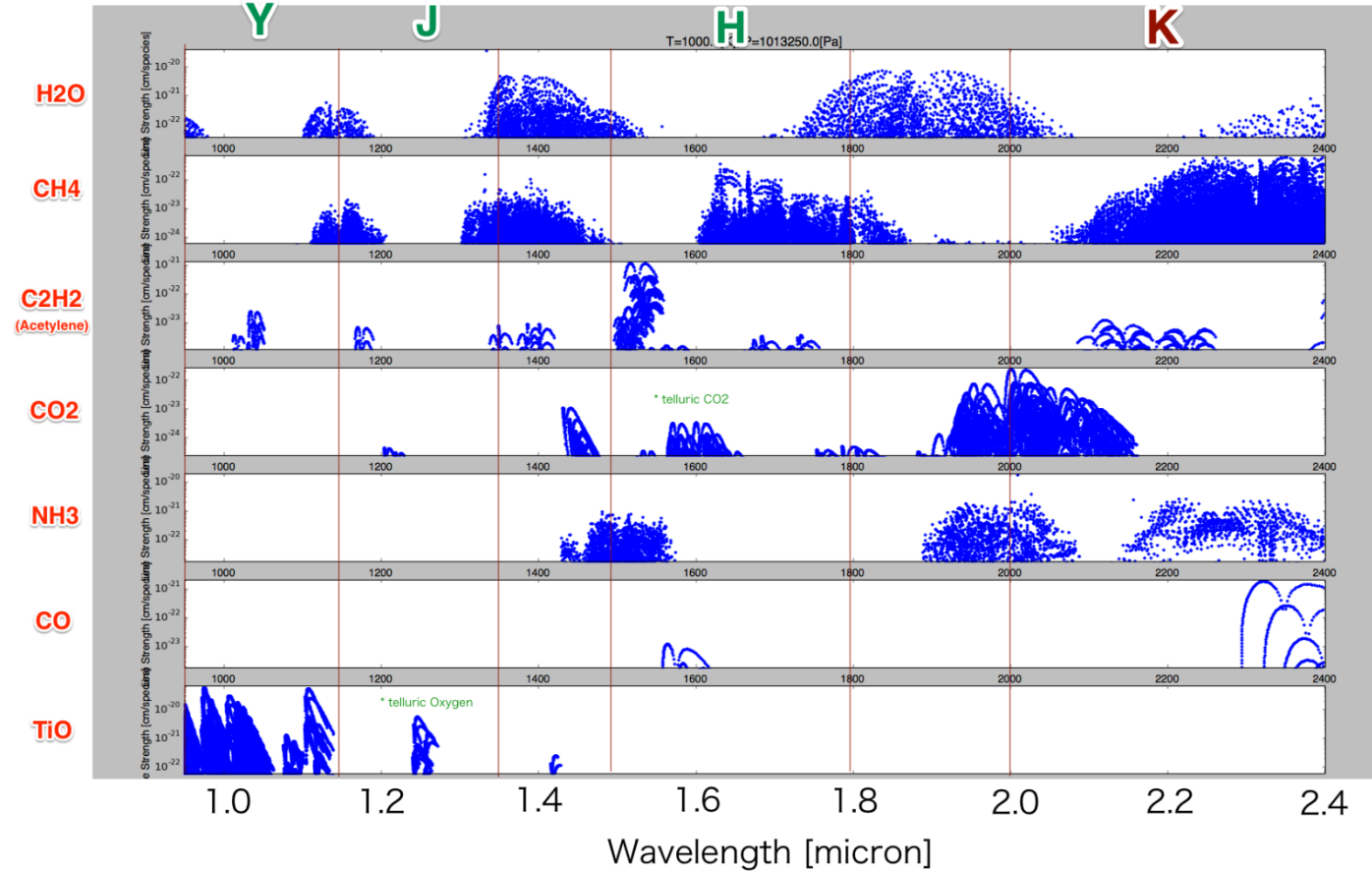
↓ evolution



H₂O rich atmosphere

HCN is also available in H band

Use KPIC!

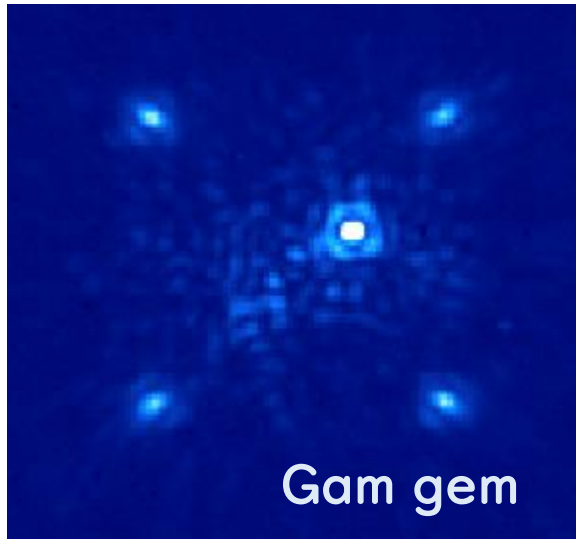




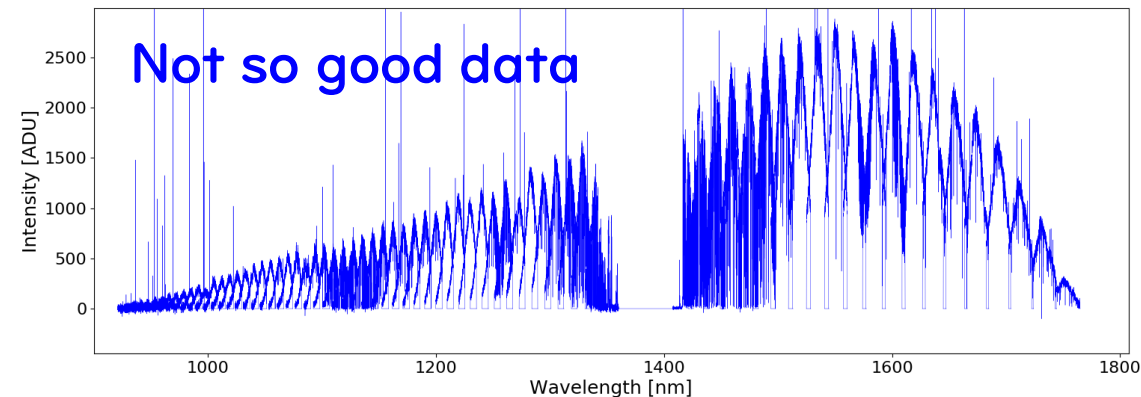
Science case #2: HRS-C for cool directly imaged planets



Ready for targets with low-contrast whose image is visible at raw level (HRS-NC).
But, planet characterization (HRS-C) is much more challenging.



A coronagraph suppresses $>50\%$ of speckle from a host star?



From engineering run
in Oct 16th (2019)

We need more tests for the HRS-C mode.

1. We need to inject the light based on positions computed from the orbit,
2. Which coronagraph is the best for the HRS-C observation?
3. Telluric removal technique (use another fiber close to a host star?)
4. Accuracy of line list. Always problematic for the HRS analysis. Help us!



Summary



 REACH = HCI+HRS on Subaru (y,J,H, R=100,000, northern hemisphere)
or one of the junks on Howl's moving castle

Science cases:

planet molecules, precise RV of a companion, disk science, and your science

Complementary to KPIC (K,L,M) and HiRISE (southern hemisphere)

Open use for S20B? (planned)